

[1511.05220], [1511.05975] with Seungho Gwak, Karapet Mkrtchyan, SooJong Rey

RAINBOW Valley of Colored (HS) Gravity

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100 years old

Defy Einstein Gravity?

- Super Gravity
- Higher-derivative Gravity
- Massive Gravity
- **Higher Spin Gravity**
- **Colored Gravity**



Einstein Gravity \longleftrightarrow **a single** massless spin 2

Colored Gravity:

Multiple massless spin 2
with color decoration



Q1 Isn't it wrong? No, it works in certain cases

Q2 Isn't it straightforward and boring?

There are **surprising** features

No-Go for colored gravity



- Multiple massless spin 2 $h_{\mu\nu}^I$

☞ cubic interactions

$$g_{IJK} \left(h_{\mu\rho}^I \partial^\rho h_{\nu\lambda}^J \partial^\lambda h^{K\mu\nu} + \dots \right)$$

symmetric

☞ global symmetries

$$[M_{\mu\nu}^I, M_{\rho\lambda}^J] = 4 g^{IJ}_K \eta_{[\nu[\rho} M_{\lambda]\mu}^K$$

associative

⇒ **only trivial solution**

- Color-charged massless spin 2 in **(A)dS**

☞ minimal interaction to gauge field

⇒ **violation of color gauge inv.**



Yes-Go for colored gravity

Global Symmetries

Associative
isometry

M_X

$\mathfrak{g}_i \otimes \mathfrak{g}_c$

Associative
color sym.

T_I

$$[M_X \otimes T_I, M_Y \otimes T_J] = \frac{1}{2} [M_X, M_Y] \otimes \{T_I, T_J\} + \frac{1}{2} \{M_X, M_Y\} \otimes [T_I, T_J]$$

not defined for Lie algebra

- \mathfrak{g}_c associative **color** algebra: $\mathfrak{u}(N)$
- \mathfrak{g}_i associative algebra containing **isometry**
 ➡ **contains more spectrum (e.g. HS)**

3D Chern-Simons

Colored Gravity

- CS action: $S[\mathcal{A}] = \frac{\kappa}{4\pi} \int \text{Tr} \left(\mathcal{A} \wedge d\mathcal{A} + \frac{2}{3} \mathcal{A} \wedge \mathcal{A} \wedge \mathcal{A} \right)$

- Gauge Algebra: $\mathfrak{g} = (\mathfrak{gl}_2 \oplus \mathfrak{gl}_2) \otimes \mathfrak{u}(N) \ominus \text{id} \otimes \mathbf{I}$

subtract Abelian CS

two additional gauge fields

$$[M_{ab}, M_{cd}] = 2 (\eta_{d[a} M_{b]c} + \eta_{c[b} M_{a]d}), \quad [M_{ab}, P_c] = 2 \eta_{c[b} P_{a]}, \quad [P_a, P_b] = \sigma M_{ab}$$

$$M_{ab} = \frac{1}{2} \epsilon_{ab}{}^c (J_c + \tilde{J}_c), \quad P_a = \frac{1}{2\sqrt{\sigma}} (J_a - \tilde{J}_a)$$

$$\begin{aligned} \sigma &= +1 \text{ for AdS}_3 \\ \sigma &= -1 \text{ for dS}_3 \end{aligned}$$

$$\text{Tr}(J_a J_b) = 2 \sqrt{\sigma} \eta_{ab}, \quad \text{Tr}(\tilde{J}_a \tilde{J}_b) = -2 \sqrt{\sigma} \eta_{ab}, \quad \text{Tr}(\mathbf{T}_I \mathbf{T}_J) = \delta_{IJ}$$

3D Chern-Simons

Colored Gravity

- CS action: $S[\mathcal{A}] = \frac{\kappa}{4\pi} \int \text{Tr} \left(\mathcal{A} \wedge d\mathcal{A} + \frac{2}{3} \mathcal{A} \wedge \mathcal{A} \wedge \mathcal{A} \right)$

☹ not tangible

☹ is it over?

☹ so what?



Let's rewrite this in **METRIC** form!

solve torsion condition only for the genuine graviton (singlet spin two)

3D Colored Gravity

👉 solve torsion condition 👉

$$S = S_{\text{CS}} + \frac{1}{16\pi G} \int d^3x \sqrt{|g|} \left[R - V(\varphi, \tilde{\varphi}) + \frac{2\sqrt{\sigma}}{N\ell} \epsilon^{\mu\nu\rho} \text{Tr} \left(\varphi_\mu^\lambda D_\nu \varphi_{\rho\lambda} - \tilde{\varphi}_\mu^\lambda D_\nu \tilde{\varphi}_{\rho\lambda} \right) \right]$$

- SU(N) CS: $S_{\text{CS}} = \frac{\kappa\sqrt{\sigma}}{2\pi} \int \left[\text{Tr} \left(\mathbf{A} \wedge d\mathbf{A} + \frac{3}{2} \mathbf{A} \wedge \mathbf{A} \wedge \mathbf{A} \right) - \text{Tr} \left(\tilde{\mathbf{A}} \wedge d\tilde{\mathbf{A}} + \frac{3}{2} \tilde{\mathbf{A}} \wedge \tilde{\mathbf{A}} \wedge \tilde{\mathbf{A}} \right) \right]$
must be quantized!

- Newton's constant: $\kappa = \frac{\ell}{4NG}$

semi-classical gravity: compatible with small CS level for large N!

3D Colored Gravity

👉 solve torsion condition 👉

$$S = S_{\text{CS}} + \frac{1}{16\pi G} \int d^3x \sqrt{|g|} \left[R - V(\varphi, \tilde{\varphi}) + \frac{2\sqrt{\sigma}}{N\ell} \epsilon^{\mu\nu\rho} \text{Tr} \left(\varphi_\mu^\lambda D_\nu \varphi_{\rho\lambda} - \tilde{\varphi}_\mu^\lambda D_\nu \tilde{\varphi}_{\rho\lambda} \right) \right]$$

- Colored spinning matter: $D_\mu \varphi_{\nu\rho} = \nabla_\mu \varphi_{\nu\rho} + [A_\mu, \varphi_{\nu\rho}]$
covariant couplings!

Potential

$$V(\varphi, \tilde{\varphi}) = -\frac{1}{N\ell^2} \text{Tr} \left[2\sigma \mathbf{I} + 4 \left(\varphi_{[\mu}^\mu \varphi_{\nu]}^\nu + \tilde{\varphi}_{[\mu}^\mu \tilde{\varphi}_{\nu]}^\nu \right) + 8\sqrt{\sigma} \left(\varphi_{[\mu}^\mu \varphi_{\nu]}^\nu \varphi_{\rho]}^\rho - \tilde{\varphi}_{[\mu}^\mu \tilde{\varphi}_{\nu]}^\nu \tilde{\varphi}_{\rho]}^\rho \right) \right] \\ - \frac{16\sigma}{N^2\ell^2} \text{Tr} \left(\varphi_{[\mu}^\nu \varphi_{\rho]}^\rho - \tilde{\varphi}_{[\mu}^\nu \tilde{\varphi}_{\rho]}^\rho \right) \text{Tr} \left(\varphi_{[\nu}^\mu \varphi_{\lambda]}^\lambda - \tilde{\varphi}_{[\nu}^\mu \tilde{\varphi}_{\lambda]}^\lambda \right) + \frac{6\sigma}{N^2\ell^2} \left[\text{Tr} \left(\varphi_{[\mu}^\mu \varphi_{\nu]}^\nu - \tilde{\varphi}_{[\mu}^\mu \tilde{\varphi}_{\nu]}^\nu \right) \right]^2$$

Matter self-interaction: \sqrt{N} times **STRONGER** than gravity!

Closer look on Potential

Potential

$$V(\varphi, \tilde{\varphi}) = -\frac{1}{N \ell^2} \text{Tr} \left[2 \sigma \mathbf{I} + 4 (\varphi_{[\mu}{}^{\mu} \varphi_{\nu]}{}^{\nu} + \tilde{\varphi}_{[\mu}{}^{\mu} \tilde{\varphi}_{\nu]}{}^{\nu}) + 8 \sqrt{\sigma} (\varphi_{[\mu}{}^{\mu} \varphi_{\nu}{}^{\nu} \varphi_{\rho]}{}^{\rho} - \tilde{\varphi}_{[\mu}{}^{\mu} \tilde{\varphi}_{\nu}{}^{\nu} \tilde{\varphi}_{\rho]}{}^{\rho}) \right] \\ - \frac{16 \sigma}{N^2 \ell^2} \text{Tr} \left(\varphi_{[\mu}{}^{\nu} \varphi_{\rho]}{}^{\rho} - \tilde{\varphi}_{[\mu}{}^{\nu} \tilde{\varphi}_{\rho]}{}^{\rho} \right) \text{Tr} \left(\varphi_{[\nu}{}^{\mu} \varphi_{\lambda]}{}^{\lambda} - \tilde{\varphi}_{[\nu}{}^{\mu} \tilde{\varphi}_{\lambda]}{}^{\lambda} \right) + \frac{6 \sigma}{N^2 \ell^2} \left[\text{Tr} (\varphi_{[\mu}{}^{\mu} \varphi_{\nu]}{}^{\nu} - \tilde{\varphi}_{[\mu}{}^{\mu} \tilde{\varphi}_{\nu]}{}^{\nu}) \right]^2$$

- reduction to parity-invariant sector

$$\chi_{\mu\nu} = \sqrt{\sigma} (\varphi_{\mu\nu} - \tilde{\varphi}_{\mu\nu}) \\ \tau_{\mu\nu} = \varphi_{\mu\nu} + \tilde{\varphi}_{\mu\nu}$$

$$V(\chi) = -\frac{2 \sigma}{N \ell^2} \text{Tr} \left(\mathbf{I} + \chi_{[\mu}{}^{\mu} \chi_{\nu]}{}^{\nu} + \chi_{[\mu}{}^{\mu} \chi_{\nu}{}^{\nu} \chi_{\rho]}{}^{\rho} \right)$$

- reduction to diffeo-invariant sector

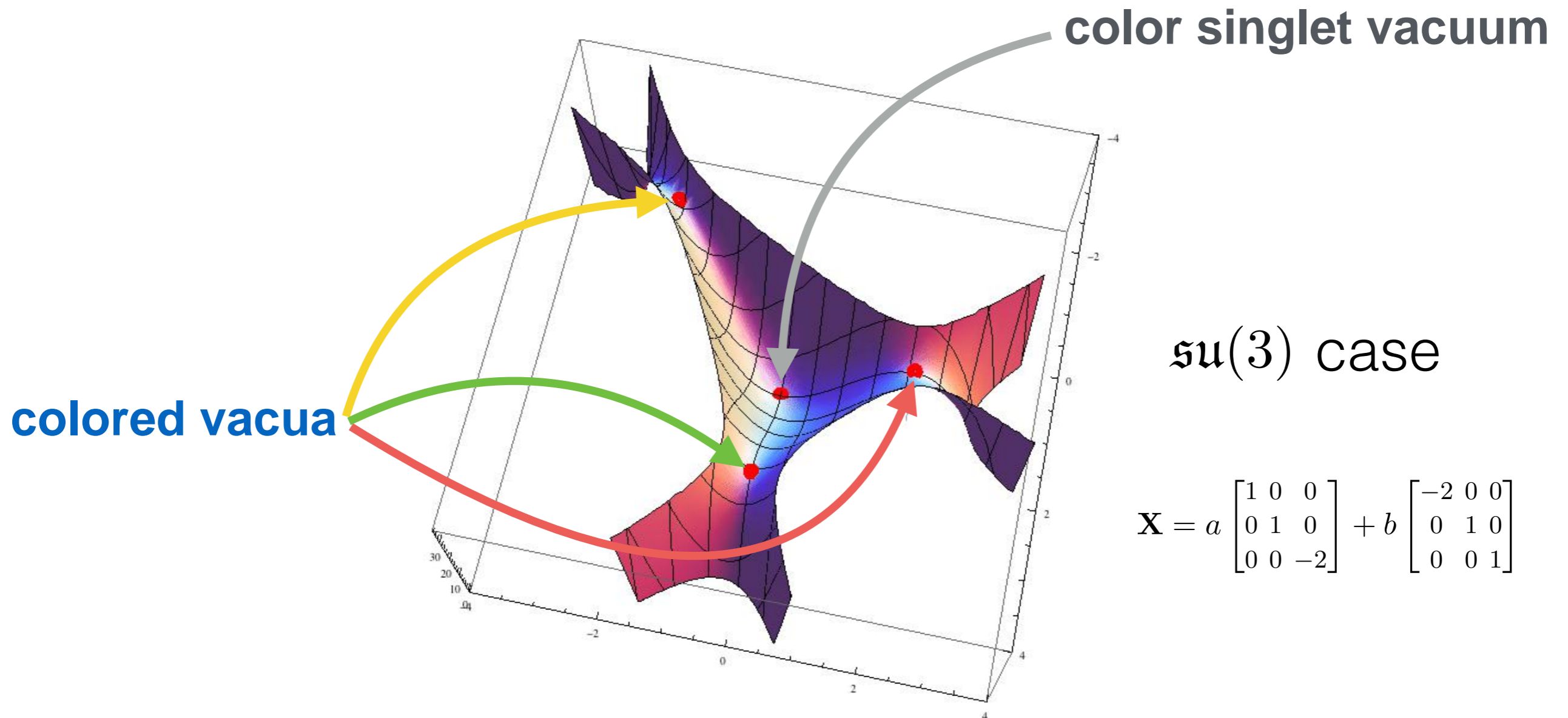
$$\chi_{\mu\nu} = g_{\mu\nu} \mathbf{X}$$

$$V(\mathbf{X}) = -\frac{2 \sigma}{N \ell^2} \text{Tr} (\mathbf{I} + 3 \mathbf{X}^2 + \mathbf{X}^3)$$

Non-trivial Potential with Many Extrema!

Rainbow Vacua

$$V(\mathbf{X}) = -\frac{2\sigma}{N\ell^2} \text{Tr}(\mathbf{I} + 3\mathbf{X}^2 + \mathbf{X}^3)$$



Rainbow Vacua



Rainbow Vacua

$$V(\mathbf{X}) = -\frac{2\sigma}{N\ell^2} \text{Tr}(\mathbf{I} + 3\mathbf{X}^2 + \mathbf{X}^3)$$

- extremum condition: $2\mathbf{X} + \mathbf{X}^2 = \frac{1}{N} \text{Tr}(2\mathbf{X} + \mathbf{X}^2) \mathbf{I}$
- solutions up to SU(N) rotation:

$$\mathbf{X} = \frac{N}{\text{Tr}(\mathbf{Z})} \mathbf{Z} - \mathbf{I} \quad \mathbf{Z}_k = \begin{bmatrix} \mathbf{I}_{(N-k) \times (N-k)} & 0 \\ 0 & -\mathbf{I}_{k \times k} \end{bmatrix}$$

Parameter

$$k = 0, 1, \dots, \left[\frac{N-1}{2}\right]$$

Symmetry Breaking

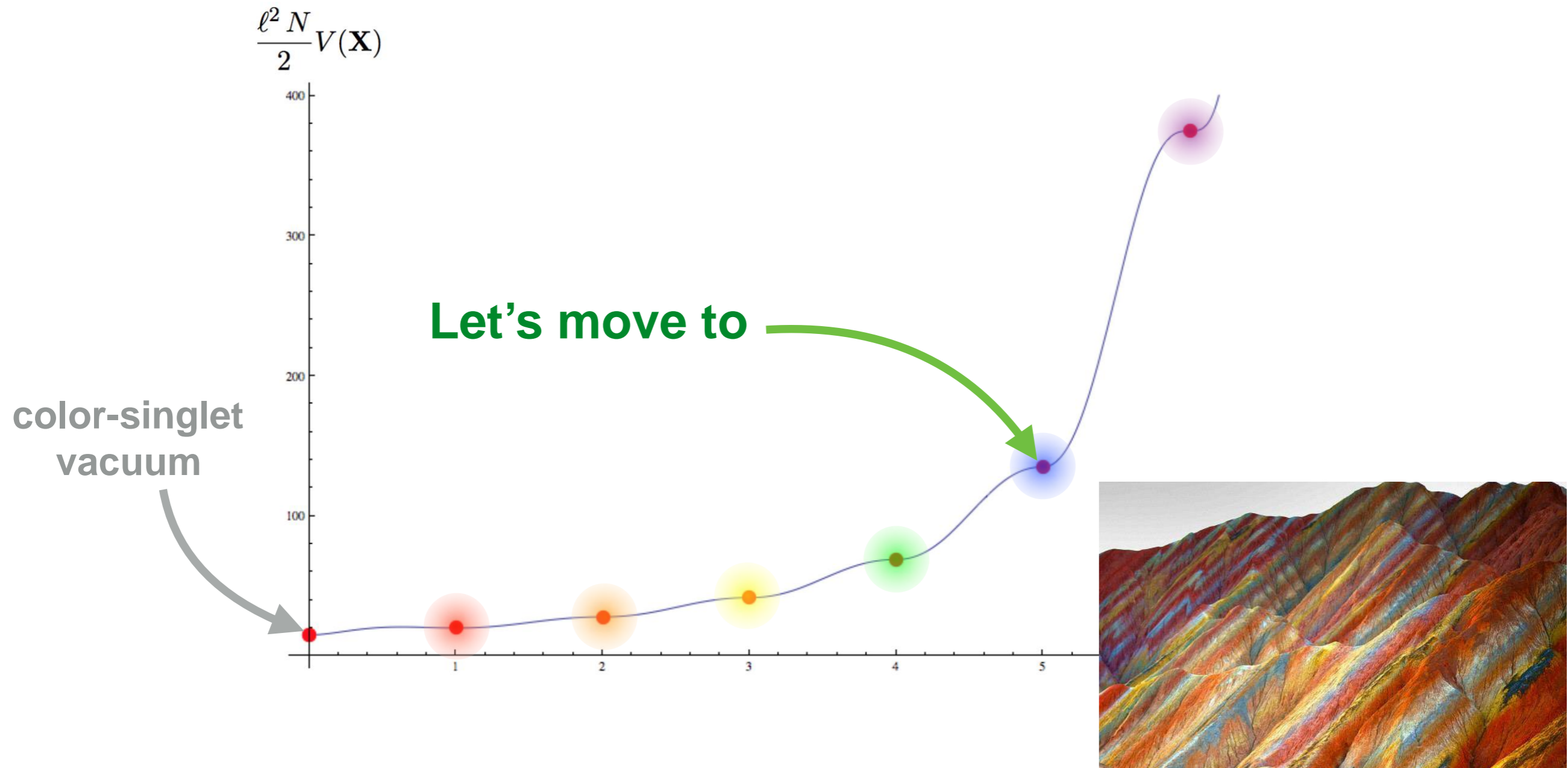
$$SU(N-k) \times SU(k) \times U(1)$$

- extremum values: $V(\mathbf{X}_k) = -\frac{2\sigma}{\ell^2} \left(\frac{N}{N-2k}\right)^2 = 2\Lambda_k$

Vacuum dependent Cosmological Constant!

Rainbow Vacua

$$V(\mathbf{X}) = -\frac{2\sigma}{N\ell^2} \text{Tr}(\mathbf{I} + 3\mathbf{X}^2 + \mathbf{X}^3)$$



Symmetry Breaking

- Diagonal parts:



- ▶ **adjoint** in $SU(N - k) \times SU(k) \times U(1)$
- ▶ still describe **massless** spin-two

- Broken-sym. part:

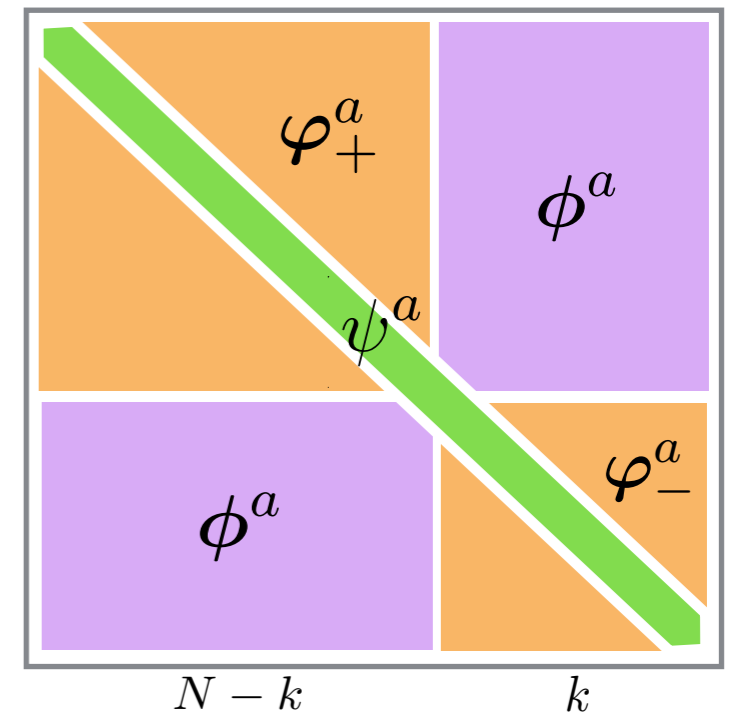


- ▶ **bi-fundamental**
- ▶ combines with (or eaten by) spin-one field

$$S_{\text{BS}}[\phi, \phi^a] = \int \phi \wedge \left(d\phi - \frac{1}{\ell} e_a \wedge \phi^a \right) - \phi_a \wedge \left(D\phi^a - \frac{1}{\ell} e^a \wedge \phi \right)$$

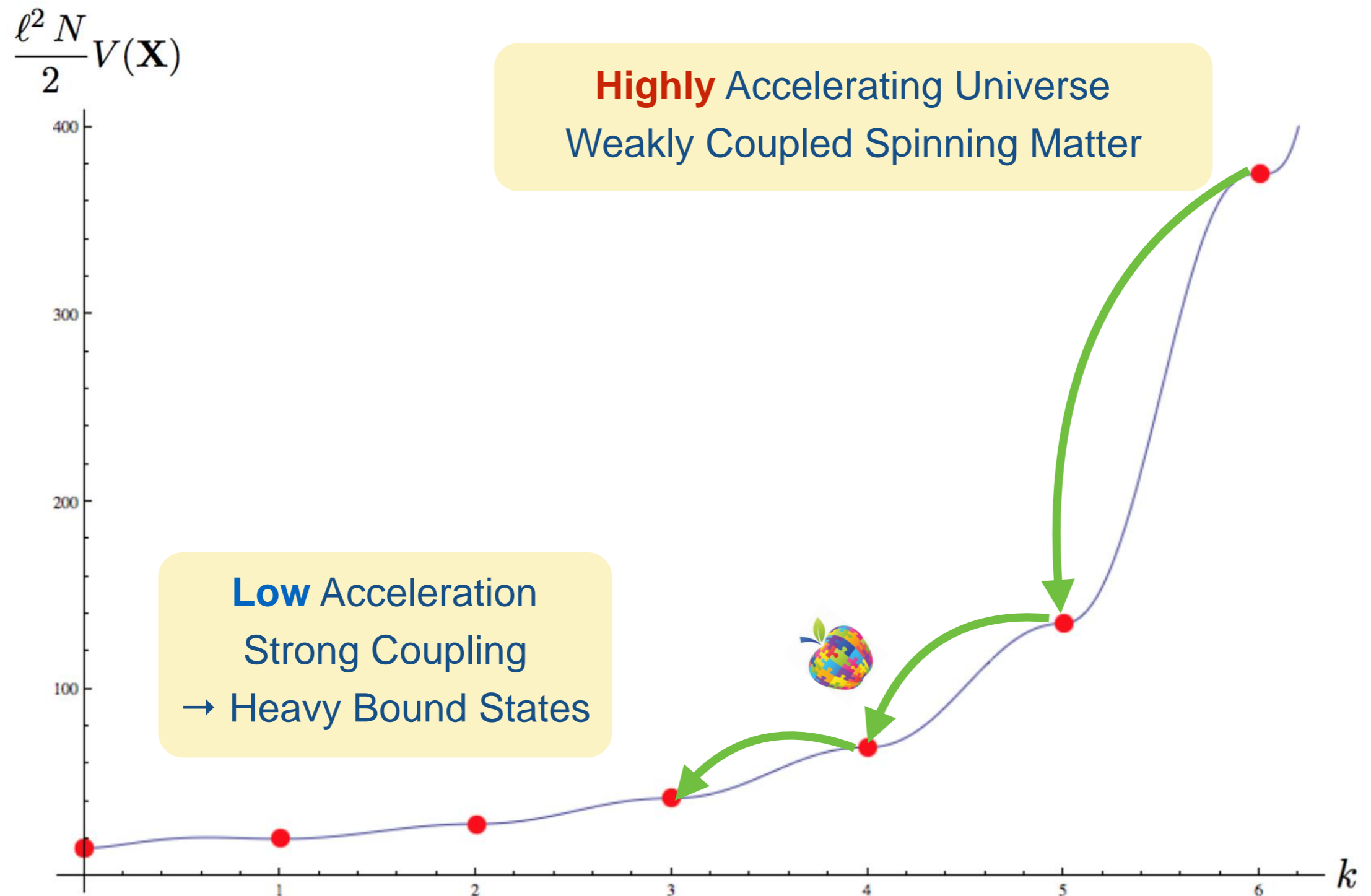
**Higgs-like
Mechanism!**

- ▶ describe **partially-massless** spin-two



- **All Weakly Interacting for Large $k \sim N/2$**

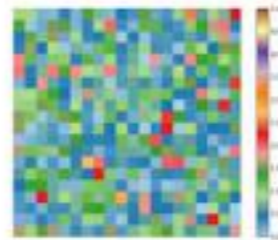
Cosmological Scenario



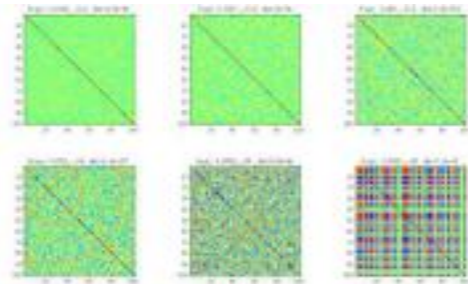
Quantum Colored Gravity

- Rainbow vacua contribution in the path integral:

Random Matrix Model



$$\mathcal{Z}_{\text{MM}} = \int d\mathbf{X} \exp [i c V(\mathbf{X})]$$



$$V(\mathbf{X}) = -\frac{2\sigma}{N\ell^2} \text{Tr} (\mathbf{I} + 3\mathbf{X}^2 + \mathbf{X}^3)$$

HS extension

- Color-decoration & Rainbow vacua extend to
 - ✓ 3D CS formulation of HS Gravity
 - ✓ Vasiliev Equations [to appear]
- Resulting spectrum after symmetry breaking
 - ▶ all the spins glue together to form an exotic one

Thank you

